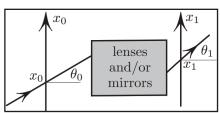
## **Telescopes and symplectic mappings**

Symplectic geometry, tracing its origins to the work of Poincaré on Hamiltonian systems, and currently a very active field, reached a high level of abstraction. Here is a simple concrete example where "symplectic" approach predicts and explains the following physical fact:

Any optical device (e.g. a telescope) which converts a parallel beam to a narrower parallel beam must necessarily magnify objects.<sup>1</sup>



**Figure 1:** We place an axis  $x_0$  before the device and another axis  $x_1$  after. The entry data  $(x_0, \alpha_0)$  of a ray determine its exit data  $(x_0, \alpha_1)$ .

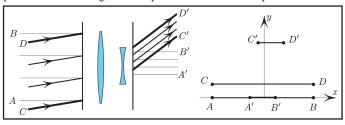
This, as I will show next, is a manifestation of the obvious fact that *if an area preserving a map squeezes in one direction, it must expand in another.* 

<sup>1</sup> To maximize simplicity, I minimize the dimension to two.

Any optical device (in two dimensions) – schematically, the black box in Figure 1 – gives rise to a map which assigns to each

ray's entry data  $(x_0, y_0)$ where  $y_0 = sin \theta_0$ , the corresponding exit data  $(x_1, y_1)$ Parallel beams, e.g. CD and C'D' in Figure 2, correspond to horizontal segments in the xy-

plane. Horizontal segments map under  $\varphi$  to



**Figure 2:** Left: Beam CD exits as beam C'D'. Right:  $\varphi$  (CD) = C'D'.

horizontal segments; moreover,  $\varphi$  shortens these segments since the device narrows parallel beams, Figure 2.

Now  $\varphi$  is area-preserving.<sup>2</sup> And since  $\varphi$  squeezes the rectangle *ABDC* (Figure

<sup>2</sup> To see why, consider the travel time  $T(x_0, x_1)$  (called the optical distance), and note that  $y_0 = -T_{x_0}(x_0, x_1)$ ,  $y_1 = T_{x_1}(x_0, x_1)$ , subscripts denoting partial differentiation. Then for a closed curve  $\gamma_0$  in the  $(x_0, y_0)$ 

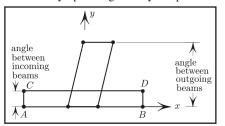
3) in the *x*-direction, it must stretch in the *y*-direction. This *y*-stretching means that the angles between parallel beams are magnified. But this is precisely what the optical magnification of objects amounts to. For example, the reason a telescope allows us to tell that a distant speck is actually a ship and not a dot is that it increases the

angle between two beams, one from the stern and the other from the bow, thus making these beams fall onto *different* "pixels" on our retina.

The proof of area-preservation

in the footnote, due to Poincaré, admits a "hands-on" palpable mechanical interpretation, as described in [1]. More on symplectic maps and lenses can be found in the

plane, parametrized  $s \in [0,1]$  we get  $0 = \int_0^1 \frac{d}{ds} T(x_0(s), x_1(s)) ds = \int_0^1 (-y_0 x_0' + y_1 x_1') ds$   $= -\int_{y_0} y dx + \int_{\varphi(y_0)} y dx.$  remarkable book [2]. And there are interesting open questions that we do not address here on the relationship between recent results of symplectic geometry to optics.



**Figure 3:** Narrowing of the beams causes widening of the angles between beams, *i.e.* the optical magnification.

## References

[1] M. Levi, Classical Mechanics with Calculus of Variations and Optimal Control: an Intuitive Introduction. AMS, March 2014, pp. 46 and 275.

[2] V. Guillemin and S. Sternberg, Symplectic Techniques in Physics. Cambridge University Press, 1984 or later editions.

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